

Notes on statistical Mechanics*

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Classical Ideal gas: canonical ensemble

Let us consider ideal monatomic gas where there is no interaction between any two atoms and no internal degrees of freedom.

We have

$$p(E)dE = \frac{\exp(-\beta E) g(E) dE}{\int_0^{\infty} \exp(-\beta E) g(E) dE} \quad (1)$$

and

$$Q_N(V, T) = \int_0^{\infty} \exp(-\beta E) g(E) dE. \quad (2)$$

Here $Q_N(V, T)$ is the partition function. Now $g(E) dE = \rho d\Gamma$, no. of microstates in the volume element $d\Gamma$. The average or expectation value of a physical quantity f is given by

$$\langle f \rangle = \int f(q, p) \rho(q, p) d^{3N}q d^{3N}p \quad (3)$$

where $\rho(q, p)$ represents the density of representative points of the system in phase space. Here $\rho(q, p) \neq \rho(t)$, because the states are equilibrium states. In fact

$$\rho(q, p) \propto \exp(-\beta H(q, p))$$

where $H(q, p)$ is the Hamiltonian of the system. Therefore

$$\langle f \rangle = \frac{\int f(q, p) \exp(-\beta H) d\omega}{\int \exp(-\beta H) d\omega} \quad (4)$$

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and

$$Q_N(V, T) = \frac{1}{N!h^{3N}} \int e^{-\beta H} d\omega \quad (5)$$

For ideal monatomic gas, no internal degrees of motion is considered. This is confined in volume V and is in equilibrium at temperature T . Because of no intermolecular interaction, the energy is wholly kinetic, and the Hamiltonian is

$$H = \sum_{i=1}^N \frac{p_i^2}{2m} \quad (6)$$

p_i s are the momenta of the gas molecules. Therefore

$$\begin{aligned} Q_N(V, T) &= \frac{1}{N!h^{3N}} \int e^{-\beta \sum_{i=1}^N \frac{p_i^2}{2m}} d^{3N}q d^{3N}p \\ &= \frac{1}{N!h^{3N}} \int \prod_{i=1}^N e^{-\frac{p_i^2}{2mkT}} d^{3N}q d^{3N}p \\ &= \frac{V^N}{N!h^{3N}} \left[\int_0^\infty e^{-\frac{p^2}{2mkT}} d^3p \right]^N \\ &= \frac{V^N}{N!h^{3N}} \left[\int_0^\infty e^{-\frac{p^2}{2mkT}} 4\pi p^2 dp \right]^N \\ &= \frac{1}{N!} \left[\frac{V}{h^3} \int_0^\infty e^{-\frac{p^2}{2mkT}} 4\pi p^2 dp \right]^N \\ &= \frac{[Q_1(V, T)]^N}{N!} \end{aligned} \quad (7)$$

$Q_1(V, T)$ in eqn. (7) may be looked upon as the single particle partition function.

Now

$$Q_1(V, T) = V \left(\frac{2\pi mkT}{h^2} \right)^{\frac{3}{2}} \quad (8)$$

using standard formula

$$\int_0^\infty e^{-Rr^2} r^{2n} dr = \frac{\Gamma(n + \frac{1}{2})}{2R^{n+\frac{1}{2}}} \quad (9)$$

Therefore,

$$Q_N(V, T) = \frac{1}{N!} \left\{ V \left(\frac{2\pi mkT}{h^2} \right)^{\frac{3N}{2}} \right\}^N \quad (10)$$

and the Helmholtz's free energy

$$\begin{aligned} A &= -kT \ln Q_N(V, T) \\ &= -NkT \ln \left\{ \left(\frac{V}{N} \right) \left(\frac{2\pi mkT}{h^2} \right)^{\frac{3N}{2}} + 1 \right\} \end{aligned} \quad (11)$$

The average energy of the system is

$$\begin{aligned} E \text{ (or } U) &= -\frac{\partial}{\partial \beta} [\ln Q_N(V, T)] \\ &= \frac{3}{2} NkT \end{aligned} \quad (12)$$

The pressure exerted by an ideal classical gas is

$$p = -\left(\frac{\partial A}{\partial V} \right)_{T, N} = \frac{NkT}{V} = nkT \quad (13)$$

Now the entropy of the system,

$$\begin{aligned} S &= -\left(\frac{\partial A}{\partial T} \right)_{V, N} \\ &= Nk \left[\ln \left\{ \left(\frac{V}{N} \right) \left(\frac{2\pi mkT}{h^2} \right)^{\frac{3}{2}} \right\} + \frac{5}{2} \right] \end{aligned} \quad (14)$$

Eqn. (14) is just the **Seckur-Tetrode** equation as already obtained from micro-canonical ensemble.

References

- [1] Statistical Mechanics: R. K. Pathria