

Q. Fraunhofer double slit diffraction pattern is observed in the focal plane of a lens of focal length 0.5 m. The wavelength of incident light is 500 nm. The distance between two maxima adjacent to the maximum of zero order is 5 mm and the fourth order maximum is missing. Find the width of each slit and the distance between their centres.

Ans: As for missing of 4th order maximum we have,

$$\frac{a+b}{a} = 4$$

$$\Rightarrow b = 3a$$

Now, for 1st order maximum,

$$(a+b) \sin \theta_1 = \lambda$$

$$\text{as } \theta_1 \text{ is small } \rightarrow \sin \theta_1 \approx \theta_1 \Rightarrow \theta_1 \approx \frac{\lambda}{a+b} = \frac{\lambda}{4a}$$

\(\therefore\) Distance between two maxima adjacent to the maximum of zero order is,

$$x = f \times 2\theta_1 = \frac{2f\lambda}{4a}$$

$$\Rightarrow a = \frac{f\lambda}{2x} = \frac{0.5 \times 500 \times 10^{-9}}{2 \times 5 \times 10^{-3}} \text{ m} = 0.025 \text{ mm}$$

$$\text{So } b = 3a = 0.075 \text{ mm.}$$

Q. A diffraction grating 2 cm wide is just able to resolve sodium D-lines in second order. Find the number of rulings per mm.

Ans! we know

$$\frac{\lambda}{d\lambda} \approx N \cdot m$$

$$\text{Now, } m=2, \lambda = \frac{589 + 589.6}{2} \text{ nm} = 589.3 \text{ nm}$$

$$d\lambda = 0.6 \text{ nm}$$

$$\text{Therefore, } N = \frac{\lambda}{m \cdot d\lambda} = \frac{589.3}{2 \times 0.6} = 491$$

$$\therefore \text{No. of rulings per mm} = \frac{\text{Total no. of lines}}{\text{width of grating in mm}} = \frac{491}{20} = 24.5$$

Q: A parallel beam of light of wavelength 500 nm is incident normally on a narrow slit of width 0.2 mm. The Fraunhofer diffraction is observed on a screen which is placed at the focal plane of a convex lens of focal length 20 cm. Calculate the approximate distance between first two maxima.

Ans: We know that the 1st two maxima occur at,

(From Theory) \rightarrow

$$\alpha_1 (\approx 3 \cdot \frac{\pi}{2}) = 1.43 \pi$$

$$\alpha_2 (\approx 5 \cdot \frac{\pi}{2}) = 2.46 \pi$$

So, corresponding angles of diffraction are given by.

$$\left. \begin{aligned} a \sin \theta_1 &= 1.43 \lambda \\ a \sin \theta_2 &= 2.46 \lambda \end{aligned} \right\}$$

$$\theta_1 = \sin^{-1} \left[\frac{1.43 \times 500 \times 10^{-9}}{0.2 \times 10^{-3}} \right] = 3.57 \times 10^{-3} \text{ rad}$$

$$\text{Similarly } \theta_2 = \sin^{-1} \left[\dots \right] = 6.15 \times 10^{-3} \text{ rad}$$

\therefore The linear distance between the first two maxima is $f \times (\theta_2 - \theta_1) = 0.516 \text{ mm}$

Q: The objective of Telescope has a diameter 2.54 m. Assuming the mean wavelength of incident light to be 550 nm. Calculate the least angular separation of two stars which can be resolved by it.

Ans: The least resolvable angle = $\theta = \frac{1.22 \lambda}{D}$ radian

Now, $D = 2.54 \text{ m}$, $\lambda = 550 \times 10^{-9} \text{ m}$,

$$\theta = \frac{1.22 \times 550 \times 10^{-9}}{2.54} = 2.642 \times 10^{-7} \text{ radian}$$

$$\theta = \left(2.642 \times 10^{-7} \times \frac{180}{\pi} \times 3600 \right)$$

$$= (\quad)$$

Q. Consider a grating of width 5 cm with three slits of width $a = 0.001$ mm separated by a distance of 0.002 mm. How many orders would be visible at $\lambda = 550$ nm? Calculate the width of principle maxima. Would there be any missing orders?

Ans:

For Grating $(a+b) \sin \theta = m\lambda$

here, $a = 0.001$ mm

$b = 0.002$ mm

$\lambda = 550 \times 10^{-9}$ mm $= 550 \times 10^{-6}$ mm

Now, for $\theta = 90^\circ$

$m = \frac{a+b}{\lambda} = \frac{0.003}{550 \times 10^{-6}} = 5.45$

So, maximum No. of observable orders = 5

According to Resolving power of a Grating
 Angular width of principle maxima = $2d\theta$

$2d\theta = \frac{2\lambda}{N(a+b) \cos \theta}$

$= \frac{2 \times 550 \times 10^{-9}}{5 \times 10^{-2}}$

$2d\theta = 2.2 \times 10^{-5}$ rad.

For maximum

$\cos \theta \approx 1$

$N(a+b) = \text{width of grating}$
 $= 5 \times 10^{-2}$ m

Now,

$\frac{a+b}{a} = \frac{3}{1}$

This indicates that 3rd, 6th, 9th etc order maxima will be missing.

Q. An oil immersion microscope just resolves 1st. rulings of a grating having $\frac{3900 \text{ lines per mm}}{400 \text{ nm}}$ when a light of wavelength 400 nm is employed. Find the Numerical Aperture (N.A.) (2)

Ans: So, the distance between two rulings = $\frac{1}{3900} \text{ mm}$.

thus,
$$\frac{1}{3900} \text{ mm} = \frac{1.22\lambda}{2(\text{N.A.})}$$

$$\text{N.A.} = 0.95$$

Q. The refractive indices of a glass for wavelengths 656.3 nm and 527.0 nm are respectively 1.6545 and 1.6635 . Calculate the length of the base of a 60° prism of this glass which can just resolve sodium lines of wavelengths 589 nm & 589.6 nm .

Ans: The change of r.i. = dn
 " " of wavelength = $d\lambda$

Now,
$$t \frac{dn}{d\lambda} = \frac{\lambda}{d\lambda}$$

$$\frac{dn}{d\lambda} = \frac{1.6635 - 1.6545}{(656.3 - 527.0) \times 10^{-9}} = 69605 \text{ m}^{-1}$$

$$\lambda = \frac{589 + 589.6}{2} \text{ nm} = 589.3 \text{ nm}$$

$$d\lambda = 0.6 \text{ nm}$$

$$t = \frac{\lambda/d\lambda}{dn/d\lambda} = \frac{589.3}{0.6} \times \frac{1}{69605} \approx 1.4 \text{ cm}$$

Assignment 5X4 = 20

Q. A convex lens of focal length 40 cm. is employed to focus the Fraunhofer diffraction pattern of a single slit of 0.3 mm width. Calculate the linear distance of first order dark band from the central band. take the wavelength of light $\lambda = 589 \text{ nm}$. [4]

H.T.

Q. How many orders would be visible, if the wavelength of incident light is 589 nm and the number of lines in the grating is 104 per mm? [4]

H.T.

Q. Sodium light of wavelengths 589.0 nm and 589.6 nm are made incident normally on a grating having 500 lines per mm. Calculate the angular dispersion of these lines in the spectrum of first order. [4]

H.T.

Q. Calculate the least width that a grating must have to resolve two components of sodium D lines (having wavelengths 589 nm and 589.6 nm) in the second order. the number of lines per mm of the grating is 80. [4]

H.T.

Q. Two stars situated at a distance of $9.5 \times 10^{12} \text{ km}$ from a telescope of diameter 20 cm are sending light of wavelength 600 nm. Find the distance of separation of the stars for what which they are just resolved. [4]